I. Origin and Age of the Universe

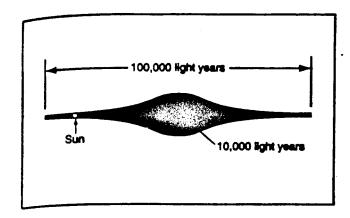
0	means everything that exists in any place – an the space,
	matter, and energy that exists in any place.
o	It is extremely and it is more than billion years old.
o	- states that all matter and energy started our
	concentrated in a small area and, after a gigantic explosion, matter began to organized into subatomic particles and atoms. The universe is still
	today.
	 Evidence of the Big Bang Theory:
	There is radiation (background radiation) from

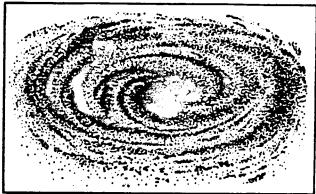
- There is _____ radiation (background radiation) from the explosion coming from all areas of the universe.
- The apparent _____ shift of most of the galaxies.
 - The ____ Effect is the shifting of wavelengths as an objects passes. ____ has a shorter wavelength and the object is coming towards you. ____ has a longer wavelength and is moving away.

J. Galaxies

0	are collections of billions of stars and various amounts of gas and
	dust held together by gravity.

- o An average galaxy has over 100 billion stars, and there are over ____ billion galaxies.
- o There are three types of galaxies based on shape:
- o Our solar system is part of the _____ Way Galaxy
 - We are part of a spiral galaxy and are located in one of the arms.







o A _______ is usually a large ball of gas held together by gravity that produces tremendous amounts of energy and shines.

o Most of the energy of stars comes from nuclear fusion.

o ______ is the combining of the nuclei of smaller elements to form the nuclei of larger elements with some of the mass being converted into energy. The sun converts Hydrogen into _____.

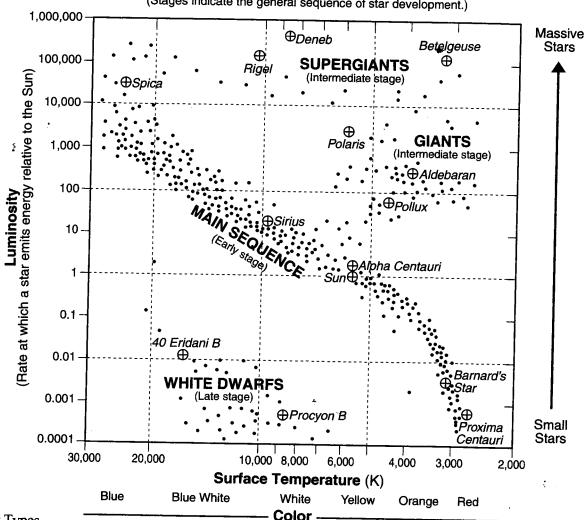
The energy of nuclear fusion is eventually radiated into space as types of electromagnetic energy.

of a star measures how bright it would be in relation to the sun if all stars were the same distance from an observer.

The colors change from red to blue as their temperatures

Characteristics of Stars

(Name in italics refers to star represented by a \oplus .) (Stages indicate the general sequence of star development.)



o Star Types

90% of stars. Can find them on the broad band on the luminosity

- o Average size stars.
- o As the size increases, the temperature
- o As the size increases, the color changes from red to blue-white.
- Our Sun is a main sequence star of yellow color.
- red, orange, and yellow giant stars are a rare type of star but are commonly seen in the night sky because of their large size.
 - o Have a low temperature because they are considered to by "dying". are very big stars that are very luminous.

o They sometimes explode in a tremendous event called a
are small (about the size of the Earth).
o They are sometimes colors besides white.
o They are hot on the surface and low in luminosity.
o They represent the luminous or shining stage of low to medium mass stars.
happens when a white dwarf cools and no longer emits electromagnetic energy.
o It is a dead star.
o it is a dead star.
The state of the second of the
I Star Origin and Evolution
Stars have an (a life cycle)
Stars originate from aloude of a life cycle)
 Stars originate from clouds of gas and dust molecules (left over from the Big
Bang).
causes the stars to clump together (forming stars)
• When the mass is high enough, nuclear fusion starts and the star begins to shine.
• Most of the life of a star is as a main sequence star.
Soveral things sould be used to the start of
• Several things could happen once the star has spent its energy (depending on size):
• Stars with masses similar to the Sun's mass become red then a
white, and then a black
Stars with masses greater than the Same
• Stars with masses greater than the Sun's mass turn into Super, and
then into a neutron star (more dense than a white dwarf).
 Stars that are extremely massive will eventually collapse to form black
0 are extreme gravity fields that allow no visible
light or any other form of energy to escape.
A SO TO STROYED TOO
- the sun and all objects that orbit the sun under its gravitational influence
• 99% of the mass in the solar system belongs to the sun.
• A is any object that orbits or revolves around another object.
• There are planets that orbit the Sun.
• An is a solid rocky and/or metallic body that orbits the sun.
O I ney have irregular shapes.
O There is a known asteroid belt between Mars and
 Are smaller than planets.
A is a body that orbits a planet or an asteroid.
There are 63 known ones in our solar system.
A is often compared to a dirty snowball.
 Are composed of solids that turn into gases easily.
 They are made of substances like ice and water.
 When comets get near the sun, their ices turn to
 When comets get near the sun, their ices turn to

 Some solids are released, forming spectacular tails visible in the Earth's sky.
are very small fragments that orbit the sun. O Most are very small.
 They leave a visual streak in the sky when they enter the atmosphere and are called
o It if actually touches the Earth's surface it is then called a
O Some meteorites have sufficient mass to create a depression in the Earth's crust
called an crater.
Evolution of the Solar System
a Sajantiata thinla that a land
years ago. billion
o Everything formed from a giant cloud of dust and gas that condensed (came
together) into a star (the Sun) and several planets.
o Planets and moons have experienced events.
events are when meteoroids, asteroids, and comets crash onto the
surfaces of planets and moons.
1 TO TI - TI OTT TI O TO
11. I I CIMOLA CIMATATALA MATERIA
1 ITIMAGU MACHICILGIGA (18) (9)
o A planet's from the sun has a major effect on its characteristics.
o Planets that are close to the sun are called planets.
• Close to the sun
Have relatively small
• Have relatively high
 Rocky surfaces have impact craters
Have few or no moons
Have no rings
Similar to Earth
• Mercury, Venus, Earth and Mars (Pluto too but it's the exception)
o Planets that are far away from the sun are called planets. • Far from the sun
• Gaseous
Have relatively largeHave relatively Small
 Have no solid surfaces (no craters) but may have a solid core.
 Have many moons
Have many rings
• Similar to Jupiter
Jupiter, Saturn, Uranus, and Neptune

Solar System Data

Celestial Object	Mean Distance from Sun (million km)	Period of Revolution (d=days) (y=years)	Period of Rotation at Equator	Eccentricity of Orbit	Equatorial Diameter (km)	Mass (Earth = 1)	Density (g/cm³)
SUN	_	_	27 d	_	1,392,000	333,000.00	1.4
MERCURY	57.9	88 d	59 d	0.206	4,879	0.06	5.4
VENUS	108.2	224.7 d	243 d	0.007	12,104	0.82	5.2
EARTH	149.6	365.26 d	23 h 56 min 4 s	0.017	12,756	1.00	5.5
MARS	227.9	687 d	24 h 37 min 23 s	0.093	6,794	0.11	3.9
JUPITER	778.4	11.9 y	9 h 50 min 30 s	0.048	142,984	317.83	1.3
SATURN	1,426.7	29.5 y	10 h 14 min	0.054	120,536	95.16	0.7
URANUS	2,871.0	84.0 y	17 h 14 min	0.047	51,118	14.54	1.3
NEPTUNE	4,498.3	164.8 y	16 h	0.009	49,528	17.15	1.8
EARTH'S MOON	149.6 (0.386 from Earth)	· 27.3 d	27.3 d	0.055	3,476	0.01	3.3

Physical Setting/Earth Science Reference Tables — 2010 Edition

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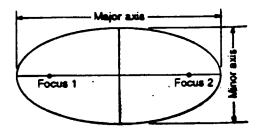
Planet Rotation

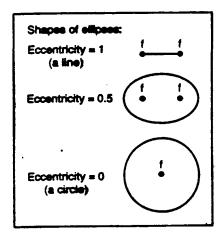
o _____ is the spinning on an imaginary axis.
o The _____ of rotation is the amount of time for a planet to spin 360 degrees. It is the length of one day on that planet.

MPlanet Revolution

0		is the movement around the sun in a path called an ellipse	€.
0	An	is the oval shape of a planets path around the sun.	

- Within the ellipse are two fixed points called
- The sun is at on foci and nothing is at the other.
- _____ is the degree of ovalness of an ellipse.
- If the eccentricity equals 1 then it would be a straight line.
- If the eccentricity equals 0 then it would be a perfect ____
- You can calculate the eccentricity of an ellipse.





Find the eccentricity of the ellipse in Figure 8-14. The dots show the locations of the two foci. You will need a centimeter scale to measure the ellipse.)

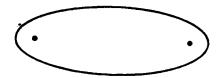
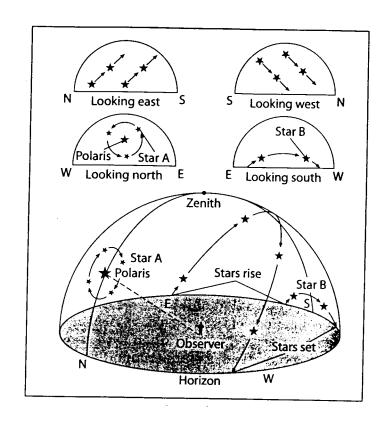


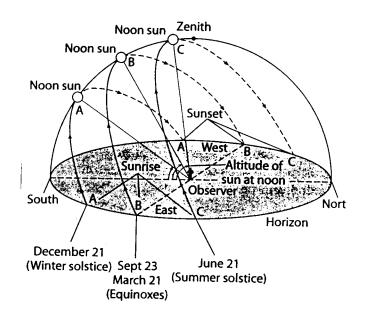
Figure 8-14.

0	The elliptical shape of planetary orbits causes the planets to vary in from the sun during revolution.
	• The Earth is closer to the sun in the (147,000,000 km)
	• The Earth is farthest from the sun in the(152,000,000 km)
	• What causes the seasons?
0	How come the Earth doesn't just fly into the sun because of the sun's gravity?
0 7	from faming in (it keeps the planet moving in sort of the same direction)
	or revolution is the amount of time it takes for a planet to revolve
	It is equal to one year on that planet.
	• The closer a planet is to the sun, the its orbit is the
	• The closer a planet is to the sun, the its orbit is, the its years are.
X Ap	parent Motions of Celestial Objects
o	is a motion that an object appears to make
	o To a person who is spinning around the room appears to move.
o	an object in the sky outside the Earth's atmosphere
	an object in the sky outside the Earth's atmosphere (the sun, moon, stars, and planets)
o	an imaginary sphere encircling the Earth on which
	all objects in the night sky appear.
0	is an uniformly curved line that is part of a circle; the path
	of the sun or a star through the sky.
О	Most celestial objects appear to move across the sky.
	Rising in the
	setting in the
0	All motion appears to move at a constant rate:
	degrees in one day.
	degrees in one hour.
	 1 degree every minutes.
O	Stars completely circle Polaris every 24 hours. They are called
	stars.
0	the movements of celestial objects over a 24 hour
	period.



	preirent moti	ions of Plemets
0 0	As seen from Earth the planets exhib	bit daily motion similar to that of stars. the planets seem to change direction in the
0	The motions of planets are not unifo	rm and are
0	The planets seem to make loops, and	back and motions.
	Like all other celestial objects the sur	IONS OF THE SUN to move through the sky
0	Its path across the sky is in the shape	of an
o	The sun's path changes both in itsseasons.	and its with the
o	Within the continental United States	
	o The sun issummer.	in the sky (72 degrees in altitude) in the
	o The sun is winter.	in the sky (28 degrees in altitude) in the
	o The noon sun is	directly overhead.
0	The sun is always at its	position in the sky at local solar noon.

	О	- the time	e at which the sun reaches its highes
	point in the	sky.	the state of the s
	0	time bas	ed on the rotation of Earth as
	reflected in	the motions of the su	ın.
0	The noon sun is only	directly	for an observer within the
	tropics.		



Teocentric Model of the Solar System

Meliocentric Model of the Solar System

Evidence of the Earth's Rotation o Thependulum.
 When the pendulum is allowed to swing freely its path will appear to
change in a predictable way.
 This is an evidence of the Earth's rotation because the pendulum
- due to would
continue to swing in the original
path if the Earth didn't rotate.
- TPI
o The effect.
 It is the tendency of all particles of matter moving at the Earth's surface to be deflected, or curve away, from a straight-line path.
o in the Northern Hemisphere
• This occurs because the Earth is rotating, and
• This occurs because the Earth is rotating, and therefore the Earth's surface is moving with respect to the path of the
particles.
Evidence of the Earth's Revolution or The changing are proof.
 The changing are proof. Because we revolve around the Sun, the two poles are tipped towards
the sun at different times of the year
o Looking at the is proof. • A is a group of stars that form a pattern and are used to help people leasts relationally in the people leasts relationally in the people leasts relationally in the people least relationally in the people relationally in t
to help people locate celestial objects.
 We can see different constellations each season because we are in
different parts of the solar system.
o The diameter of the sun appears to change throughout the year. diameter is how big it appears to be.
o Small changes in the color of stars.
• means we are moving away.
• means we are moving toward.

WTIME
is the time it takes for the Earth to rotate from solar noon
to solar on two successive days.
o is a type of time that is based on the actual motions of
the sun in the sky. (23hrs 56min)
o is exactly 24 hours.
o has been divided to make 24 hours a day for the
convenience of time keeping.
o The revolution of the moon around the Earth (as the Earth revolves around the sun) results in many common observable events. The moon revolves around the Earth in an elliptical orbit that it tilted about* from the Earth's orbit. The moon orbits the Earth once every days.
Center of mass of earth-moon system Orbit of earth-moon system around sun Orbit of mass of earth-moon system around center of earth-moon of mass of earth-moon system around center of earth-moon of mass of earth-moon system around center of earth-moon of mass of earth-moon system around center of earth-moon
Phases of the Moon
o Half of the moon is always receiving from the sun at any given time
(except for lunar eclipses).
o Since the moon revolves around the Earth, an observer on Earth sees varying
amounts of this lighted-half.
o are the changing amounts of the lighted moon as seen from

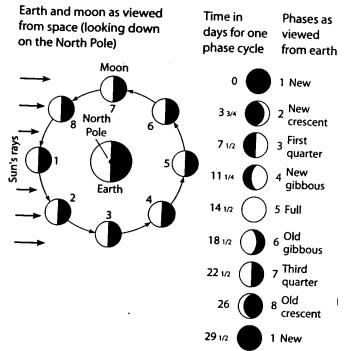
o The moon takes _____ days to revolve around us once.

o It takes ______ days for the moon to complete one complete cycle.

• An extra _____ is needed to catch up to the same spot on the Earth.

the Earth.







o _____ are the cyclic rise and fall of ocean waters

o they are caused by the gravitational attraction of the _____ and ____

o Ideally there should be 12 hours and 50 minutes between each high tide.











- o is when a celestial object partly or completely comes into the shadow of another celestial object.
- o A _____ occurs when the moon's shadow falls on a small part of the Earth and blocks out the sun.
 - It is very rare for a total eclipse of the sun (once every 200 years) and it only affects a small part of the Earth since the Moon's shadow is so small.
 - A total eclipse will only last for 7 ½ minutes.



- o A is when the Earth's shadow covers the moon.
 - It is more common, 2 total eclipses a year.
 - A total eclipse will last for over 100 minutes.
 - All the people on the dark half can see the eclipse.



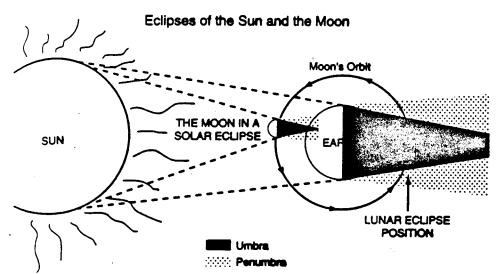


Figure 8-9. Eclipses of the sun and moon as observed from a position high above the North Pole. Eclipses of the sun fully darken a small area on the Earth. Eclipses of the moon are visible from a much larger area of the Earth.



M Seasons

o The sun's path through the sky changes with the latitude and the season because:

